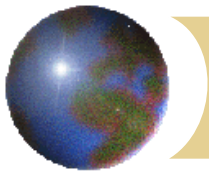


SGNoise and AGDas - tools for processing of superconducting and absolute gravity data

Vojtech Pálinkáš and Miloš Val'ko

¹Research Institute of Geodesy, Topography and Cartography, Czech Republic



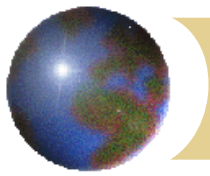


SGNoise

- ❑ Web tool primarily oriented for SG data quality control in near real time
- ❑ Written in PHP5 using the GD graphical library
- ❑ Data are processed on daily basis from raw SG data (1 sec sampling rate)
- ❑ Inspired by the previous works of **Rosat et al. (2003b, 2004, 2011, 2013)**, which established a standard feature of the noise spectrum analysis of SGs in the GGP network.
- ❑ The data quality quantification is represented by the evaluation of ambient noise level at SG stations by spectral analysis of gravity residuals and its visualization through spectrograms and probability density functions.



<http://oko.pecny.cz/grav/>



SGNoise - methodology

1 sec data, consisting of gravity and air pressure signal (**in ASCII or mseed format**)

Two basic steps based on ***Banka and Crossley (1999)***:

- ☐ computing residual gravity series
- ☐ spectral analysis of residuals

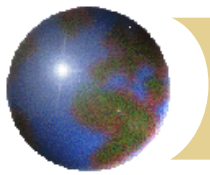
The ambient noise levels at SG stations are quantified by Power Spectral Densities of gravity residuals from 0.15 mHz (111 min) up to 61.5 mHz (16 s) .

Visualisation:

- ☐ Spectrograms
- ☐ Probability density functions (PDFs using algorithms given by ***McNamara and Buland 2004*** for seismic data)

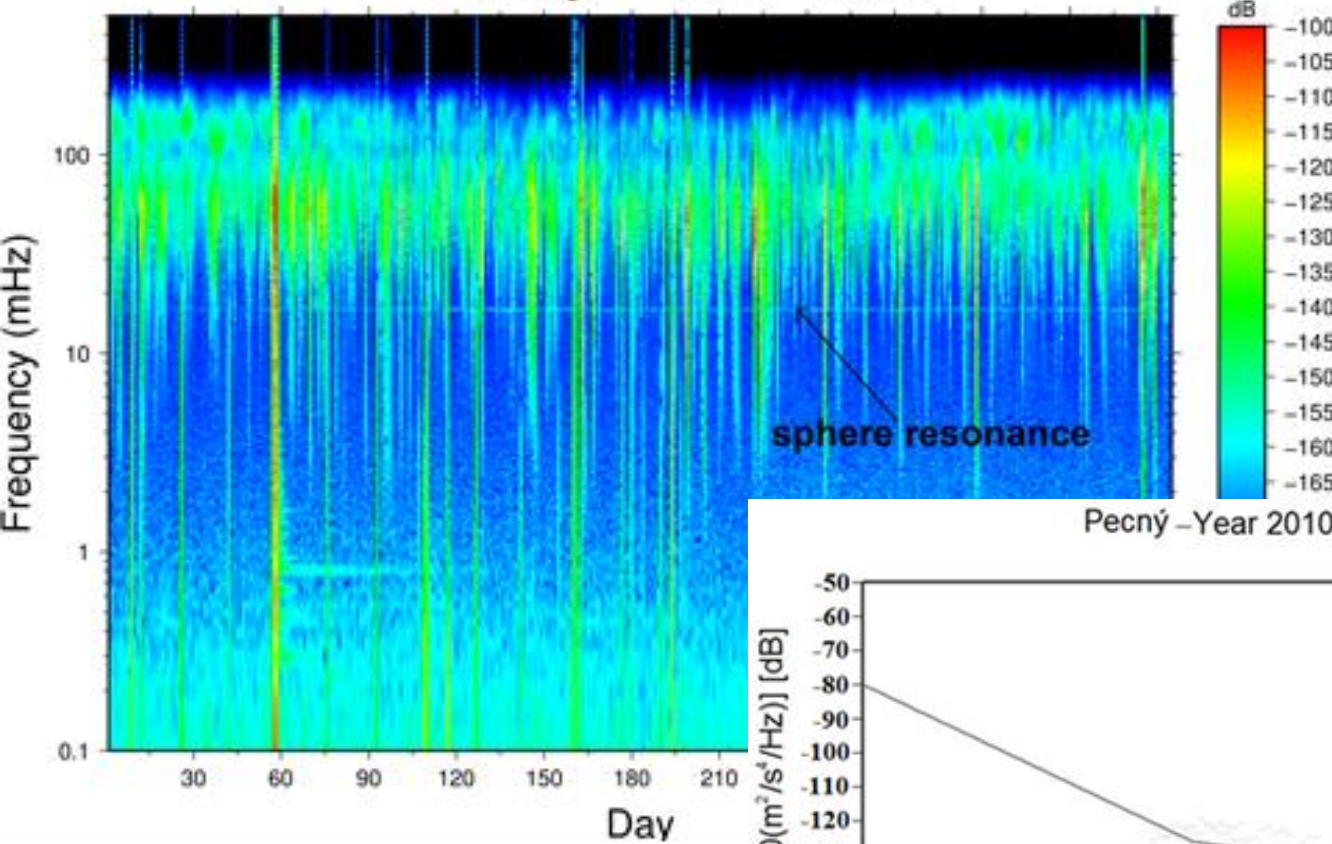
- Calibration of gravity data; Subtraction of the tides using the observed or synthetic tidal parameters; Reduction for the redistribution of atmospheric masses using local air pressure data; Subtraction of a best fitting 9th degree polynomial to eliminate the instrumental drift and any residual tidal signal.

- Windowing with a Hann window; Power Spectral Density (PSD) estimation according to *Cooley and Tukey, (1965)*; PSD smoothing using the 11-point Parzen frequency window and its expression to decibels relative to $1 \text{ (m/s}^2\text{)}^2/\text{Hz}$ through $10 \log_{10} \text{ PSD}$, where PSD is expressed in units $\text{(m/s}^2\text{)}^2/\text{Hz}$.

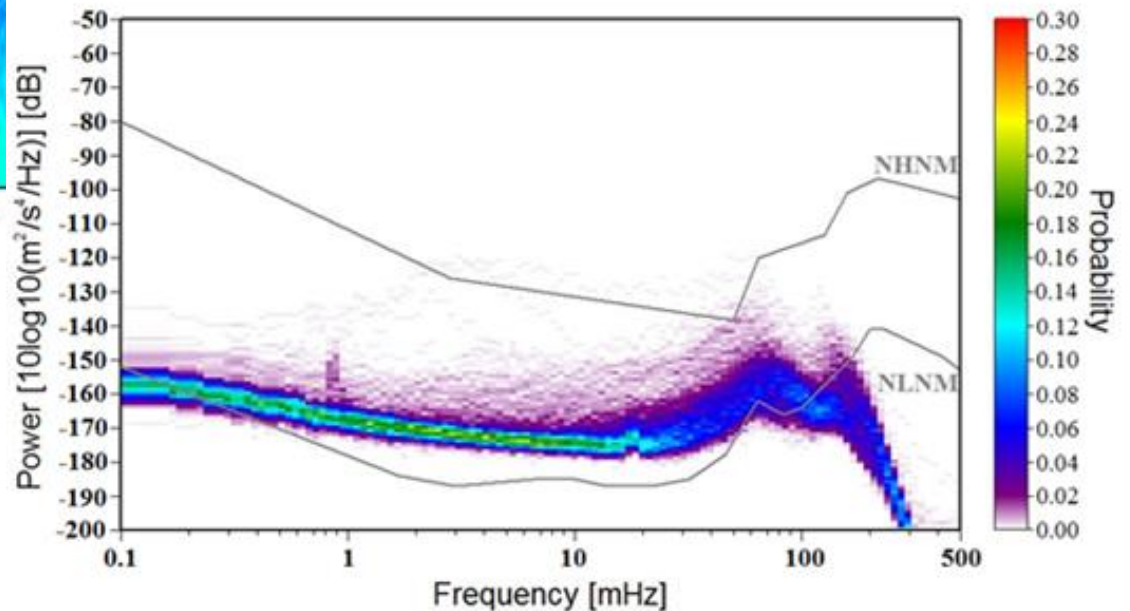


SGNoise, Visualization

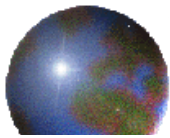
Pecný – Year 2010



Pecný – Year 2010 (730 PSD's)



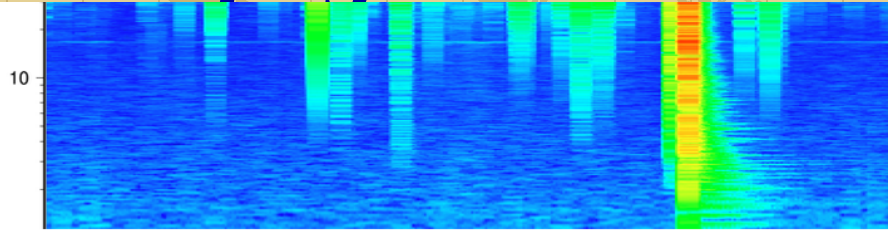
<http://oko.pecny.cz/grav/>



SGNoise, web page



Frequency (mHz)



DATA QUALITY CONTROL Station: Pecný (CZ) Year: 2014 Month: November Day: 14 Quantity: Last 30 days overview Go

MONTHLY AND YEARLY OVERVIEW Station: Pecný (CZ) Quantity: MONTHLY spectrograms Go

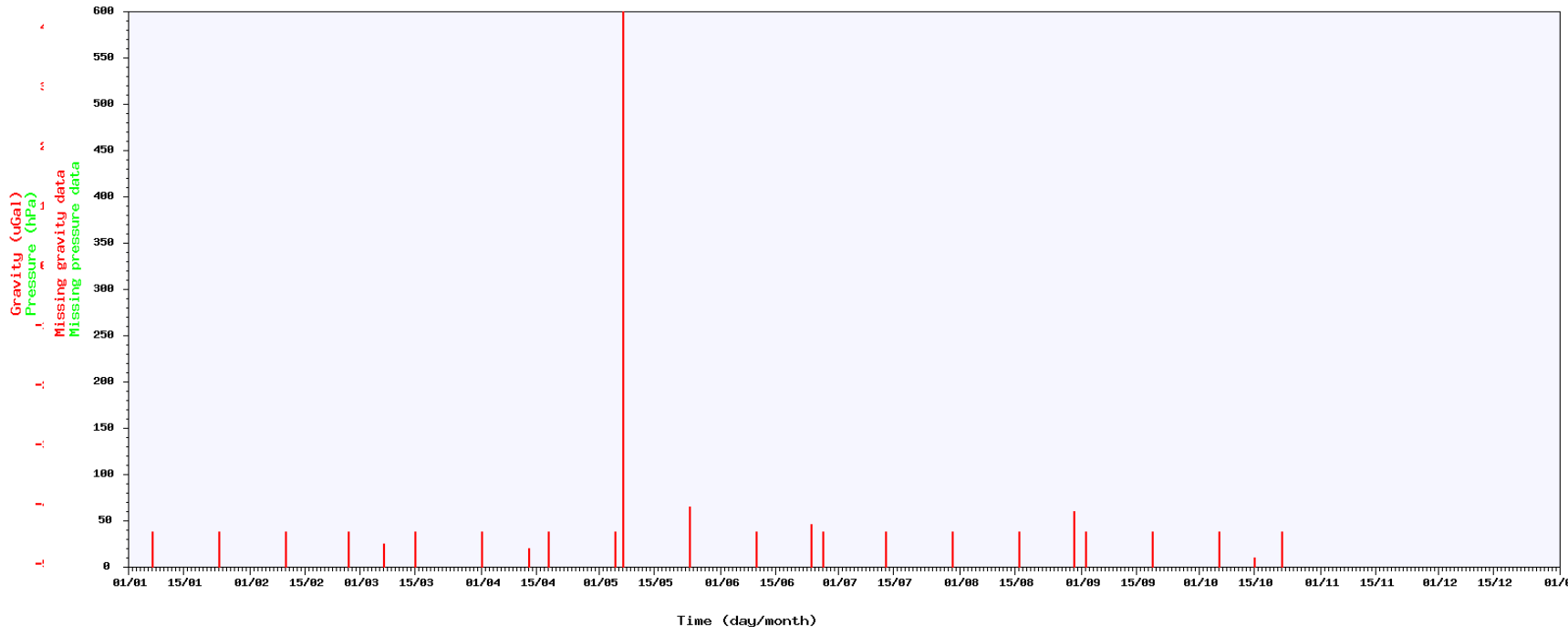
Last 30 days overview of residuals Last week spectral overview Yearly overview of residuals Yearly overview of missing data

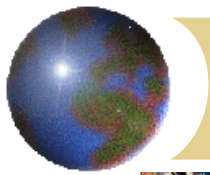
Last 30 days overview of residuals Last week spectral overview Yearly overview of residuals Yearly overview of missing data

Last 30 days overview of residuals Last week spectral overview Yearly overview of residuals Yearly overview of missing data

Last 30 days overview of residuals Last week spectral overview Yearly overview of residuals Yearly overview of missing data

Pecný, OSG-050, Yearly overview of missing data



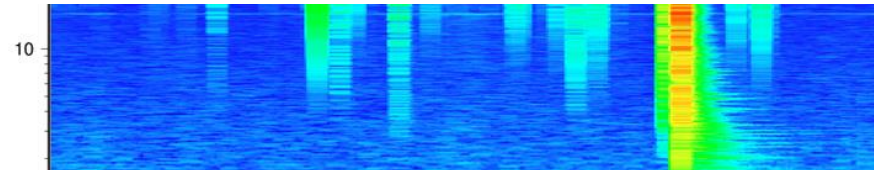


SGNoise, Daily outputs



Noise

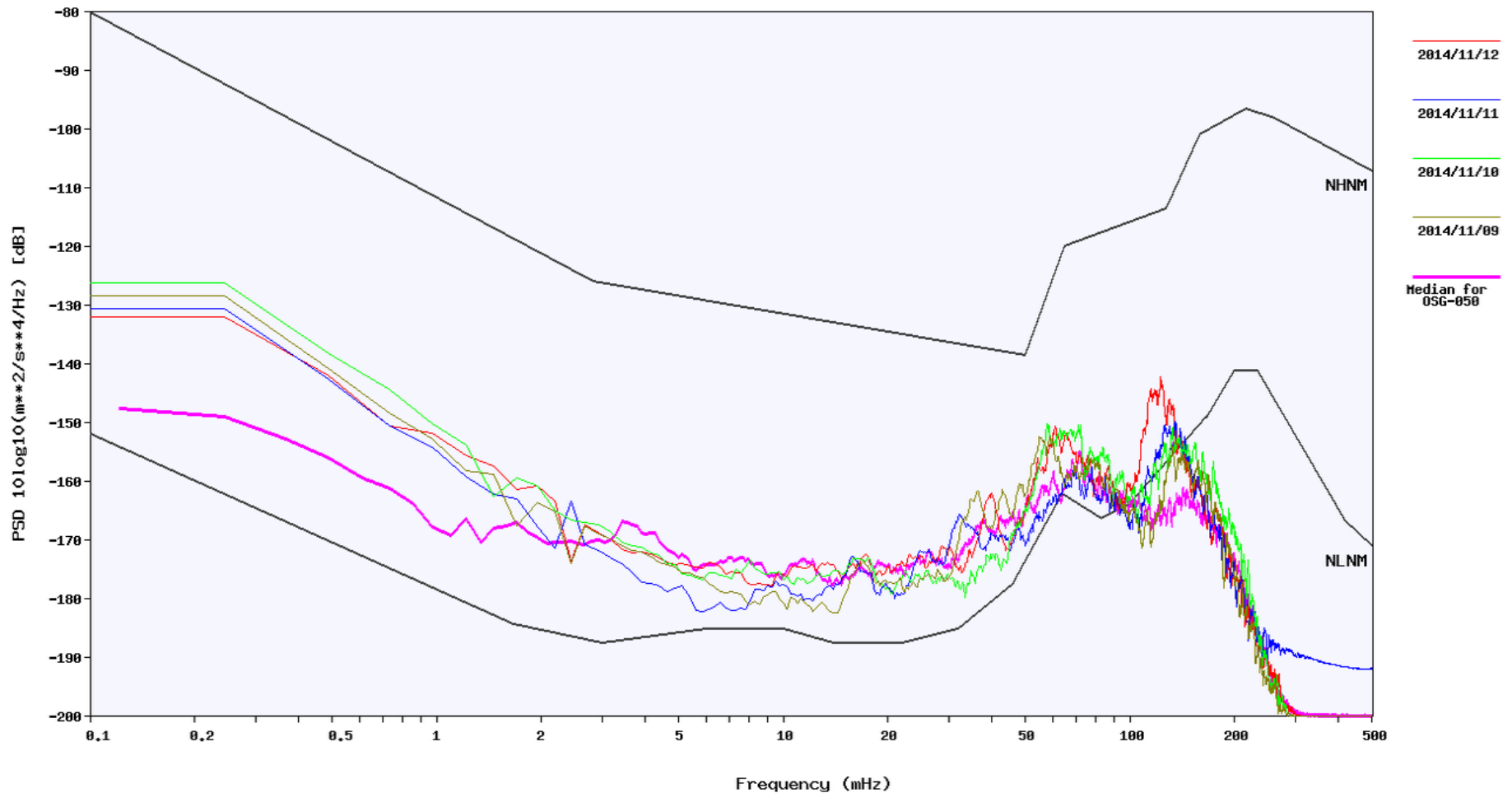
frequency (m

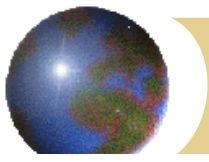


Pecny, 05G-050, residual gravity for day 2014/Nov/12

PSD for day 2014/Nov/12

PSD for days 2014/11/09 - 2014/11/12





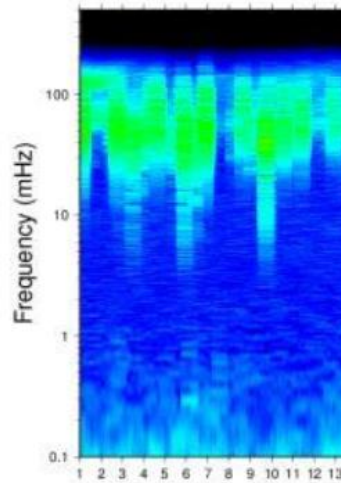
Monthly overview

MONTHLY AND YEARLY OVERVIEW

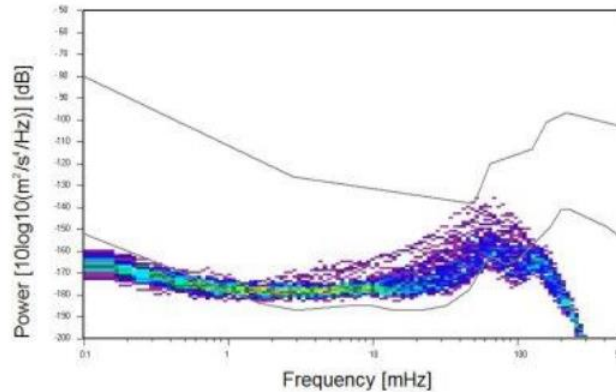
Station:

Quantity:

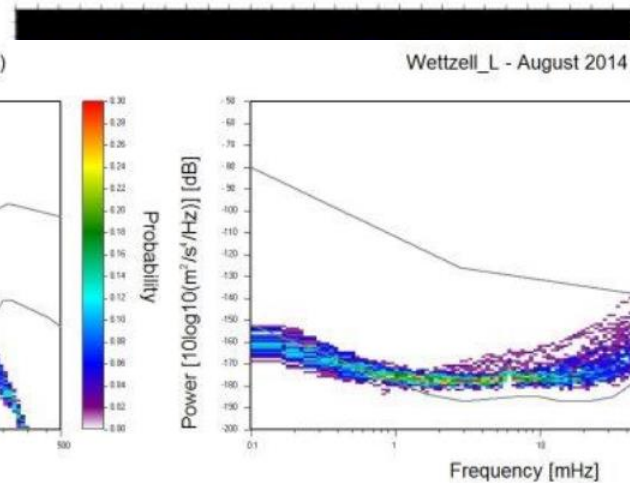
Wettzell_U – September 2014



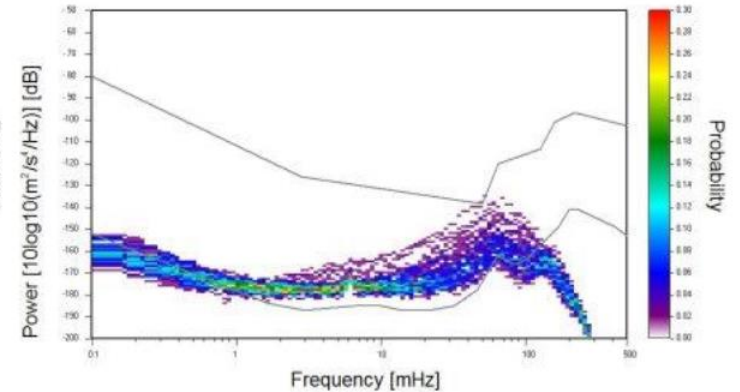
Wettzell_U - August 2014 (61 PSD's)



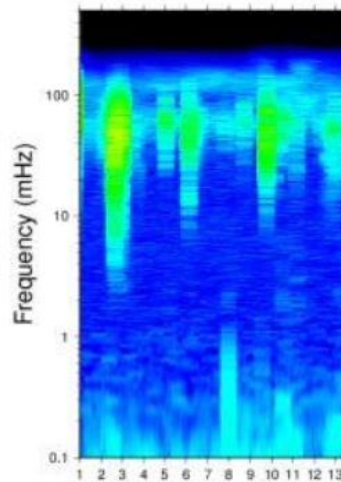
Wettzell_L – September 2014



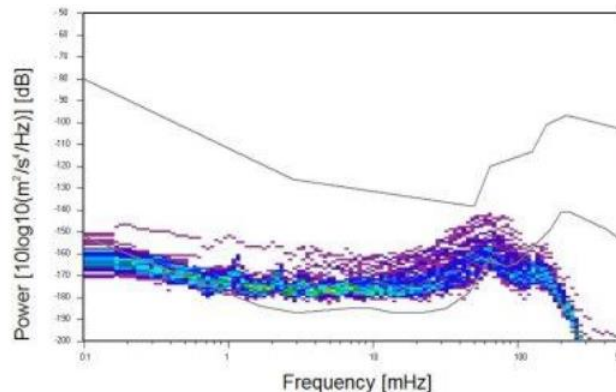
Wettzell_L - August 2014 (61 PSD's)



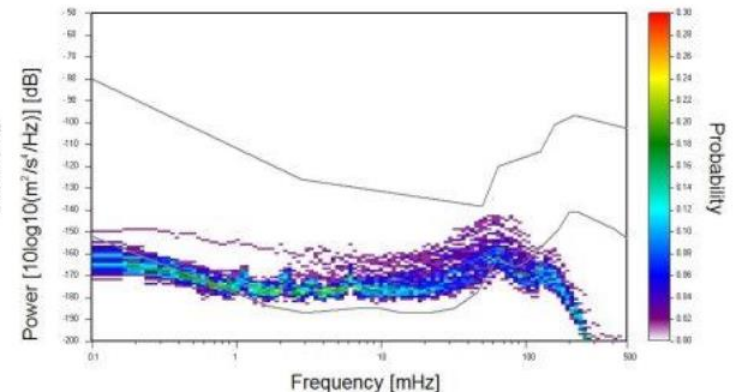
Wettzell_U



Wettzell_U July 2014 (61 PSD's)

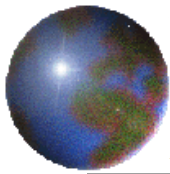


Wettzell_L - June 2014 (61 PSD's)



Day

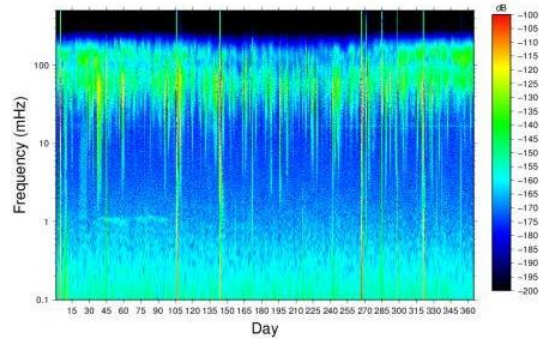
Day



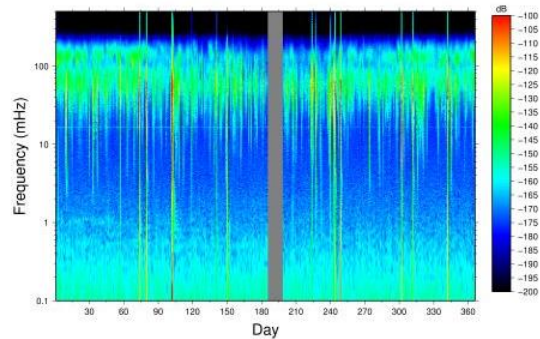
Yearly overview

MONTHLY AND YEARLY OVERVIEW Station: Quantity:

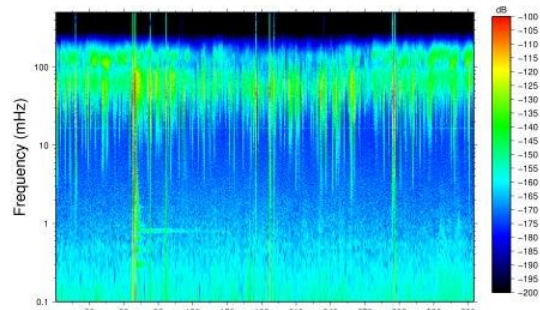
GO Pecný – Year 2013



GO Pecný – Year 2012

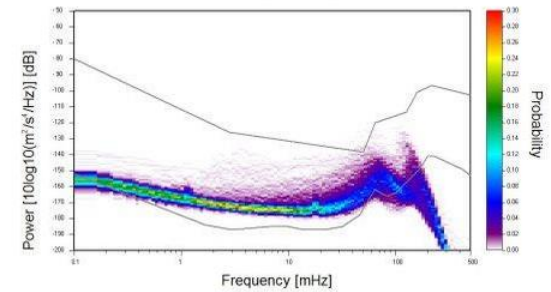


GO Pecný – Year 2011

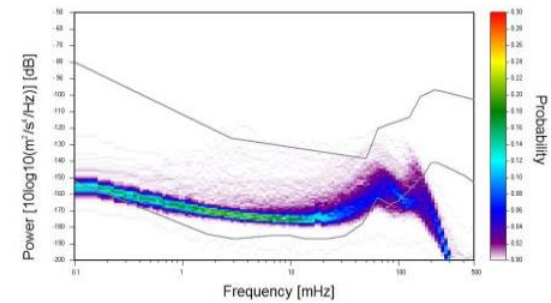


MONTHLY AND YEARLY OVERVIEW Station: Quantity:

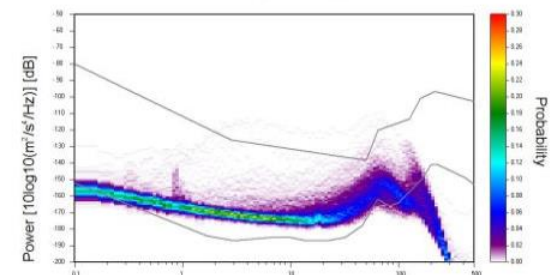
Geodetic Observatory Pecný
Year 2013 (730 PSD's)

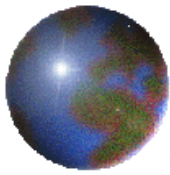


Geodetic Observatory Pecný
Year 2012 (732 PSD's)



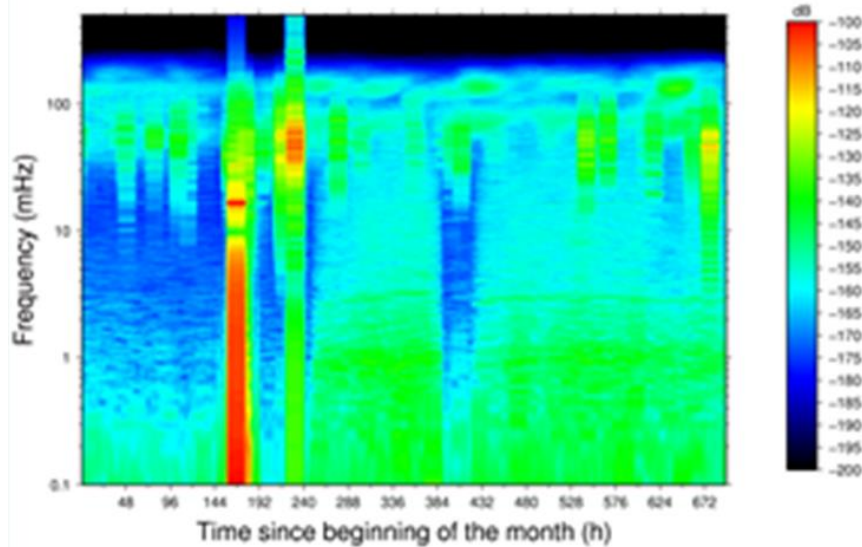
Geodetic Observatory Pecný
Year 2011 (730 PSD's)



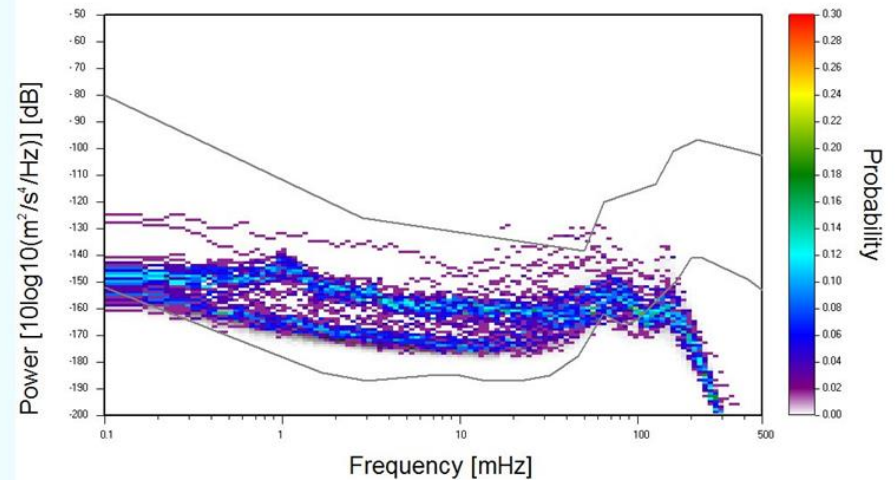


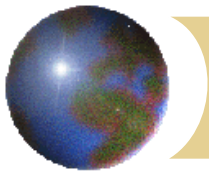
SGNoise, problem detection

GO Pecný – September 2008



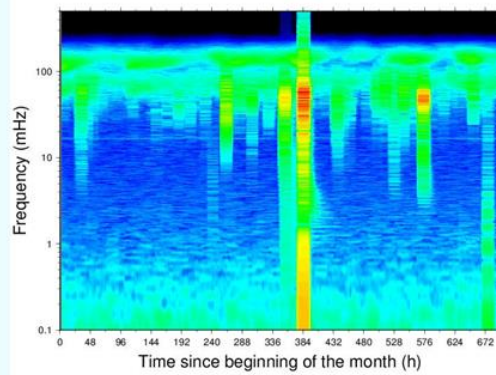
Geodetic Observatory Pecný
September 2008 (59 PSD's)



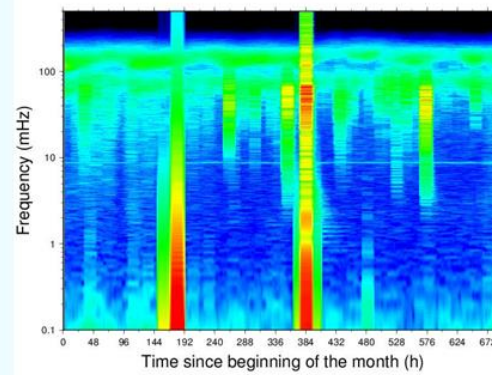


SGNoise, Comparison

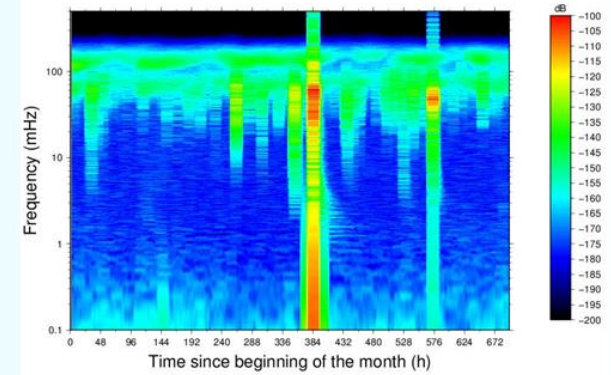
GO Pecny – November 2013



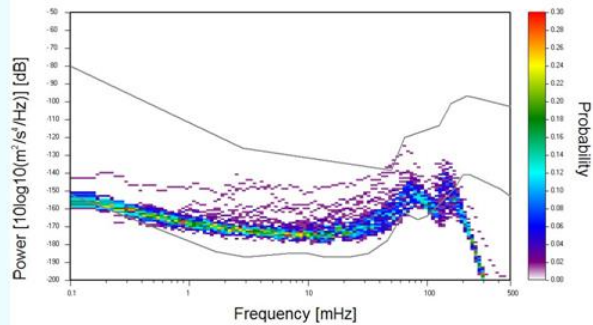
Strasbourg – November 2013



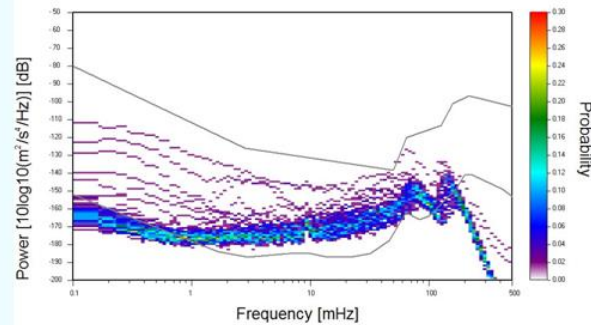
GO Wettzell – November 2013



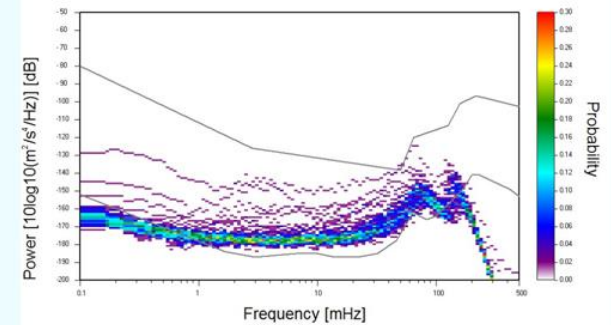
Geodetic Observatory Pecny
November 2013 (59 PSD's)

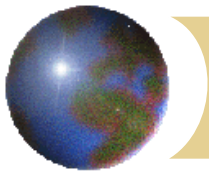


Geodetic Observatory Strasbourg
November 2013 (59 PSD's)



Geodetic Observatory Wettzell
November 2013 (59 PSD's)





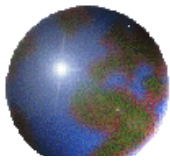
AGDAS, objective

Asolute **G**ravimeter **D**ata **A**nalysis **S**oftware
written in Matlab

Why the AGDAS has been developped?

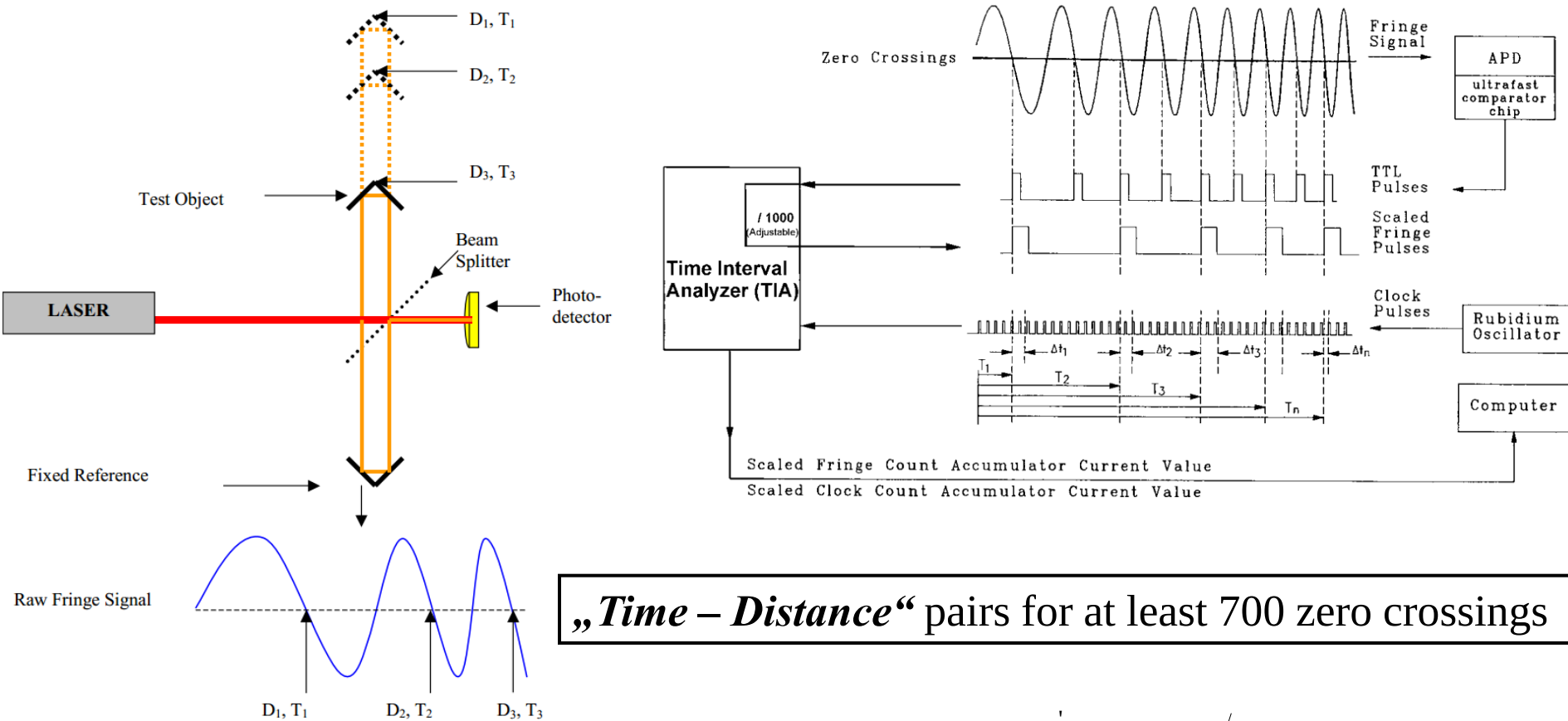
- ☐ validation of the “g” software
- ☐ accurate definition of the reference instrumental height
- ☐ complex analysis of gravity residuals in spectral and time domain
- ☐ experimental evaluation of the questionable correction from the finite speed of light.





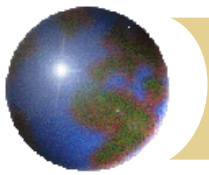
AGDAS, raw AG data

Micro-g LaCoste FG5 User's Manual



$$t'_i = t_i + z_i/c$$

$t = 0 \text{ sec} :$ $z_i = z_0 - v_0 (t'_i + \frac{\gamma}{6} t_i'^3) - g_0 (\frac{t_i'^2}{2} + \frac{\gamma}{24} t_i'^4) + a \sin(2\pi f_{\text{mod}} t_i) + b \cos(2\pi f_{\text{mod}} t_i),$



AGDAS, positioning

acceleration to the top of the drop g_{top} ($z = 0$), rather than report g_0 for time $t = 0$. The relation between these two terms is

Niebauer et al.

$$g_t = g_0 - \gamma z_0.$$

(1995)

Top of the drop

$$v = \frac{dz}{dt} = 0$$

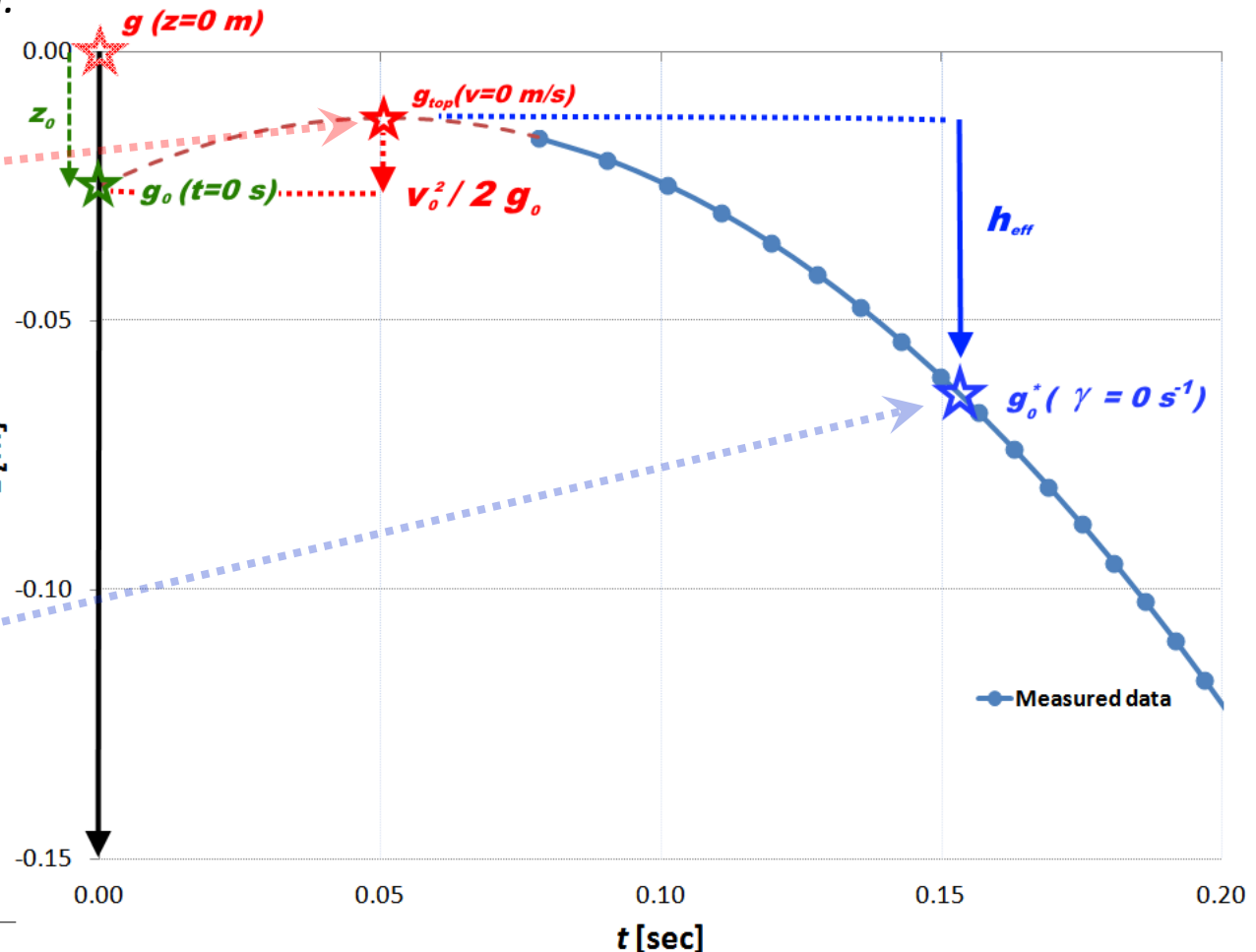
$$z = z_0 - \frac{v_0^2}{2g_0}$$

$$g_{\text{top}} = g_0 - \gamma \frac{v_0^2}{2g_0}$$

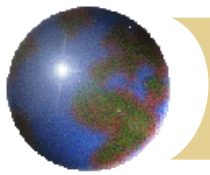
Effective position

$$z_i = z_0^* + v_0^* t_i + g_0^* \frac{t_i^2}{2}$$

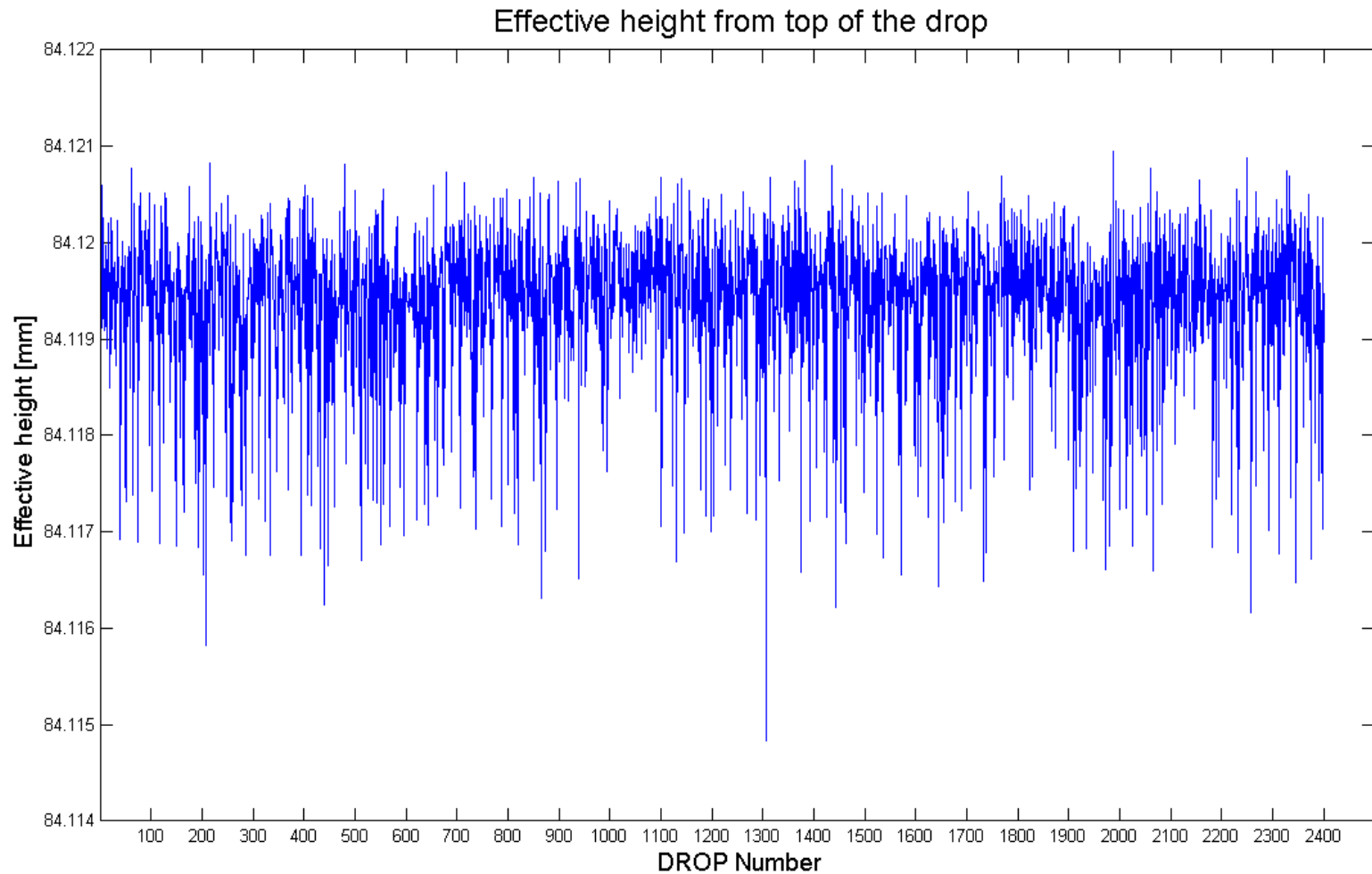
$$h_{\text{eff}} = \frac{g_0(\gamma_A) - g_0^*(\gamma = 0)}{\gamma_A} + \frac{v_0^2}{2g_0}$$

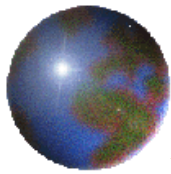


Palinkas et al. (2012)

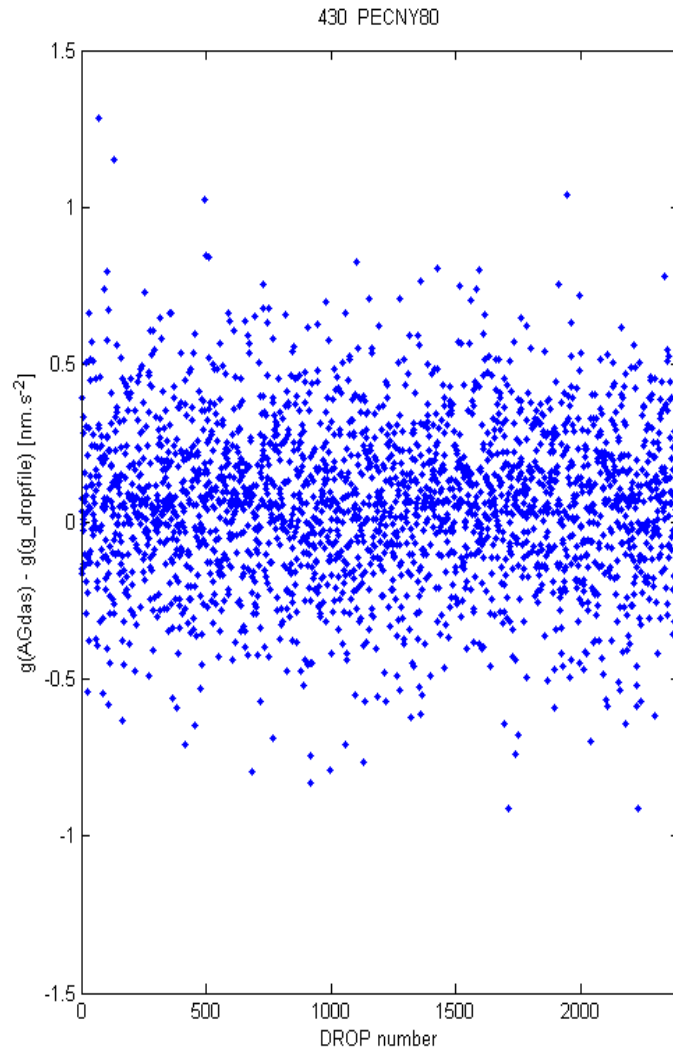


AGDAS, effective height

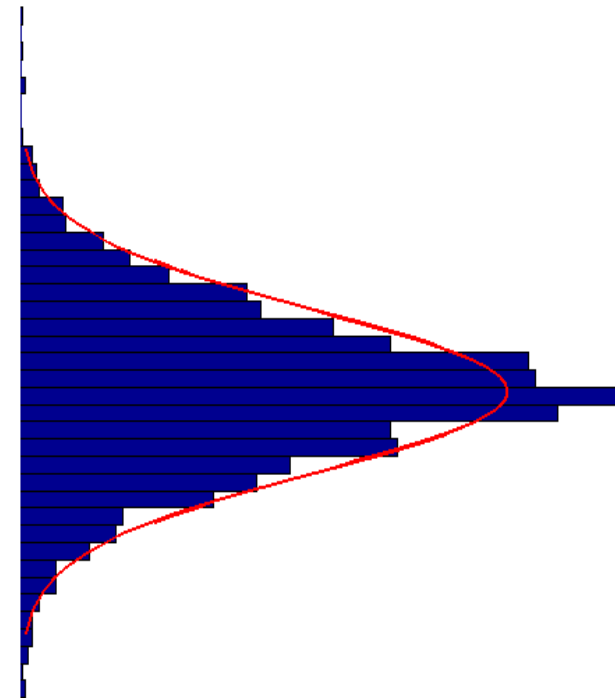




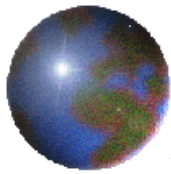
AGDAS vs. g



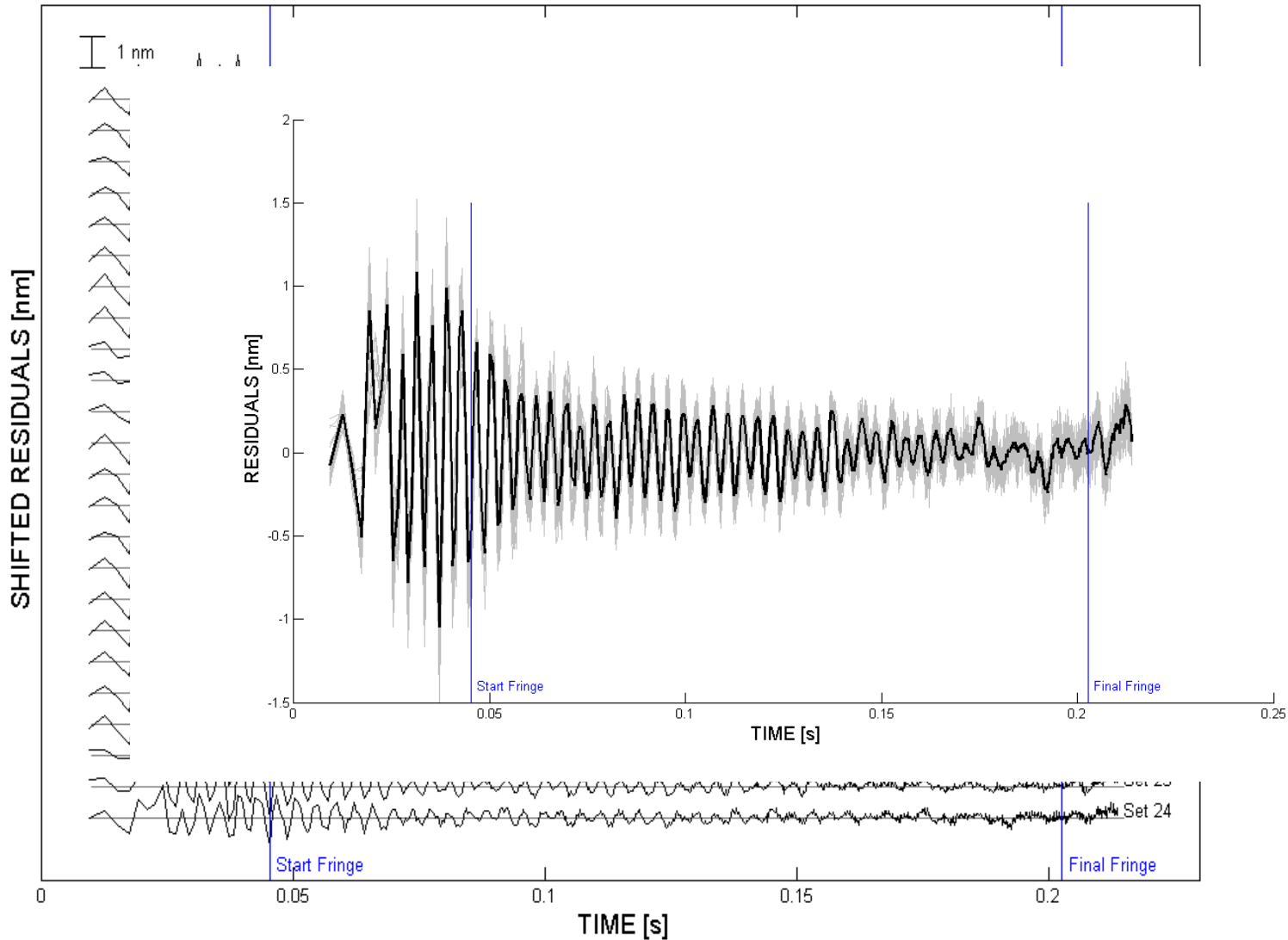
RefX0: Difference of 0.03 nm.s^{-2}

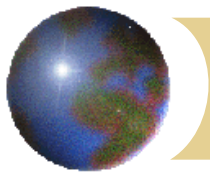


mean = 0.061 nm.s^{-2} , std = 0.0052 nm.s^{-2}

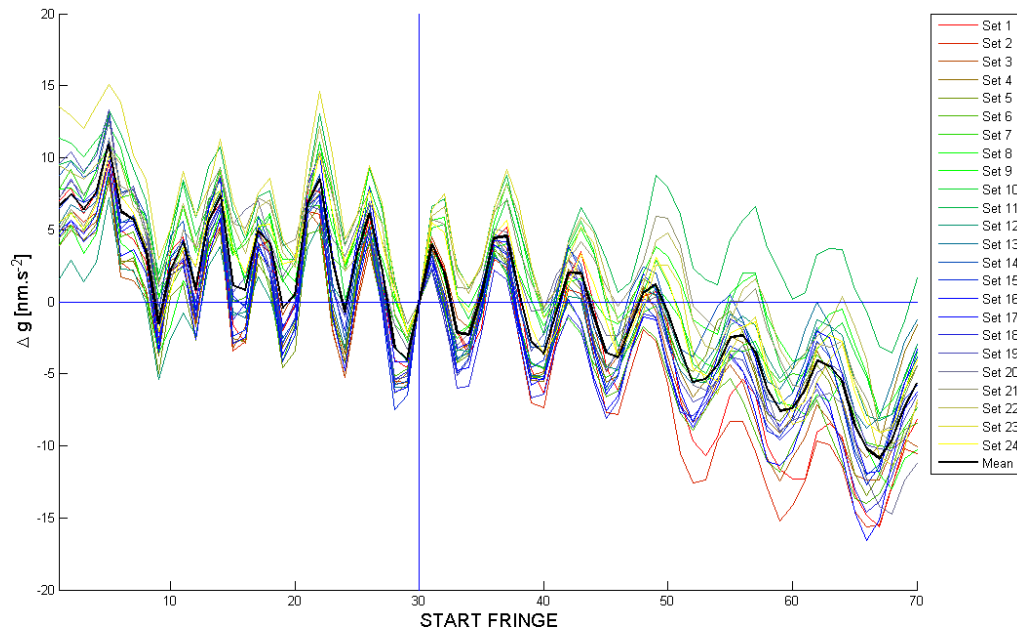


AGDAS, residuals



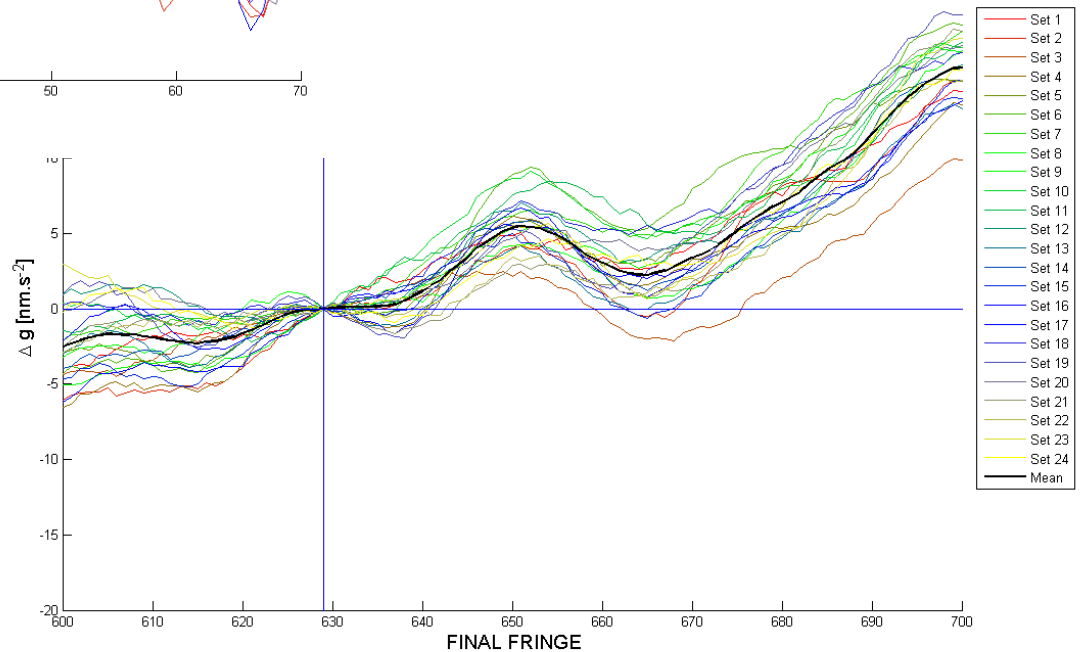


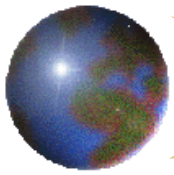
AGDAS, fringe choice



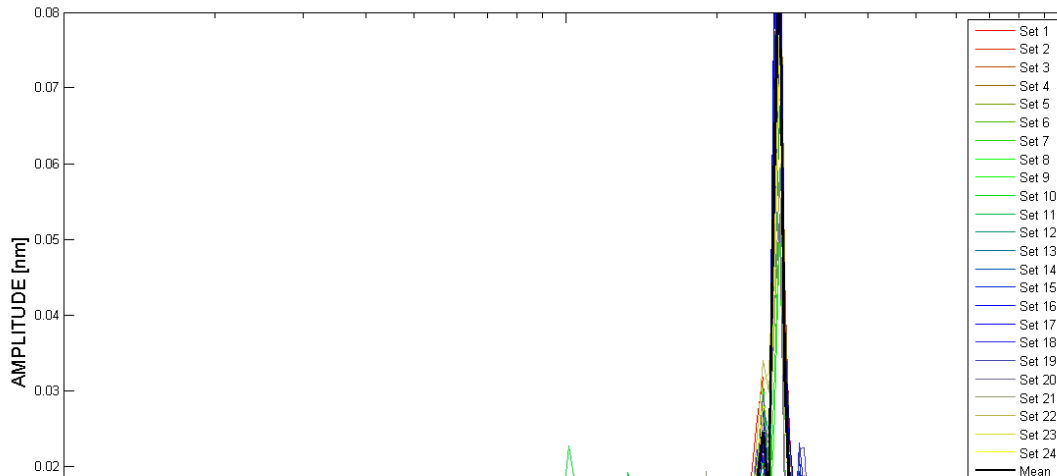
Reference solution

Fringes: 30-629





AGDAS, spectrum

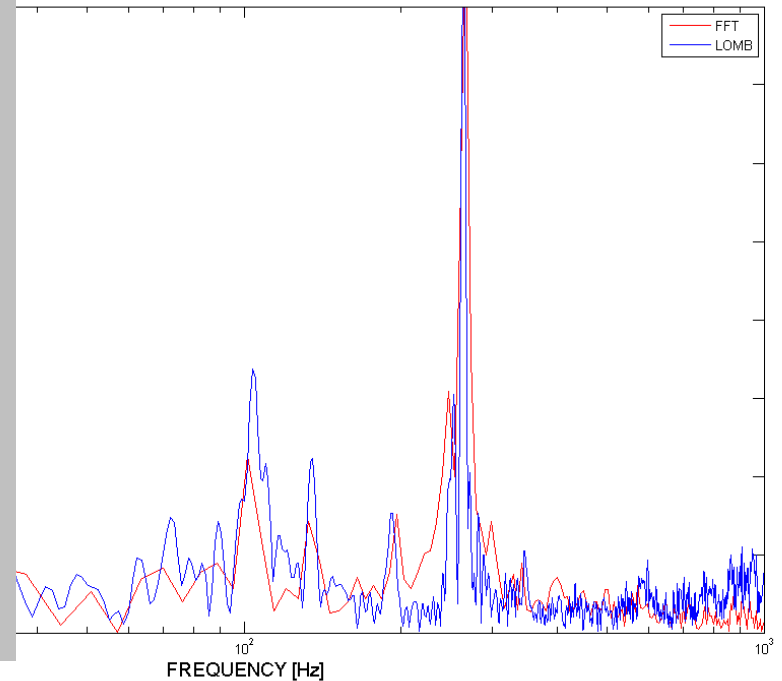
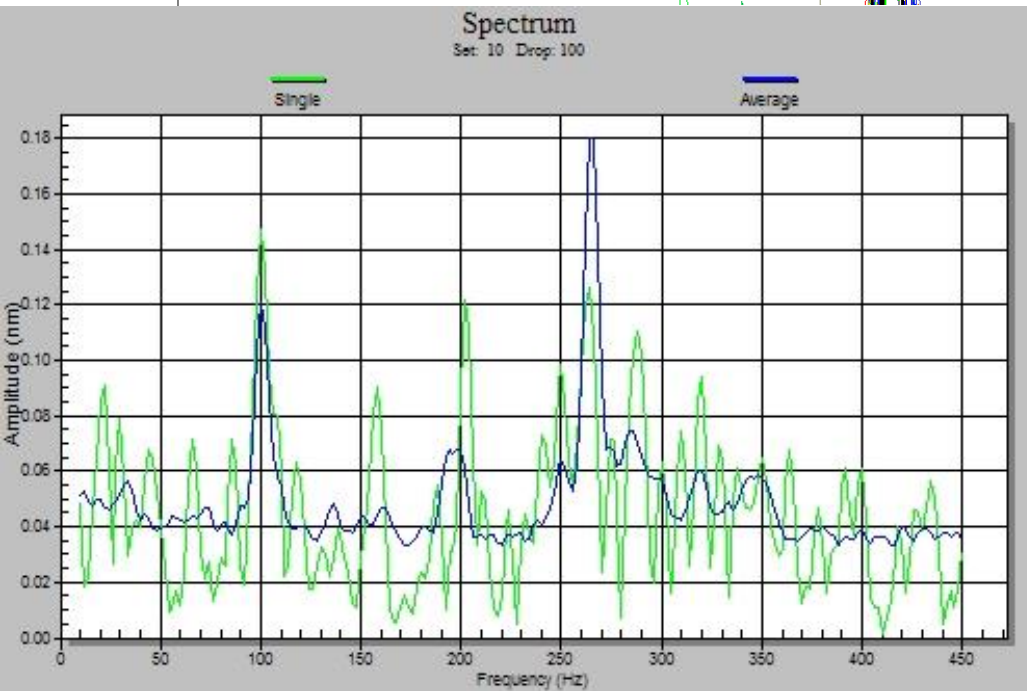


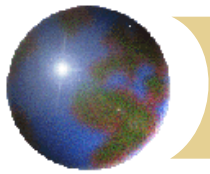
Equally spaced data

Non-equally timed data

1) Interpolation

2) Least-squares spectral analysis





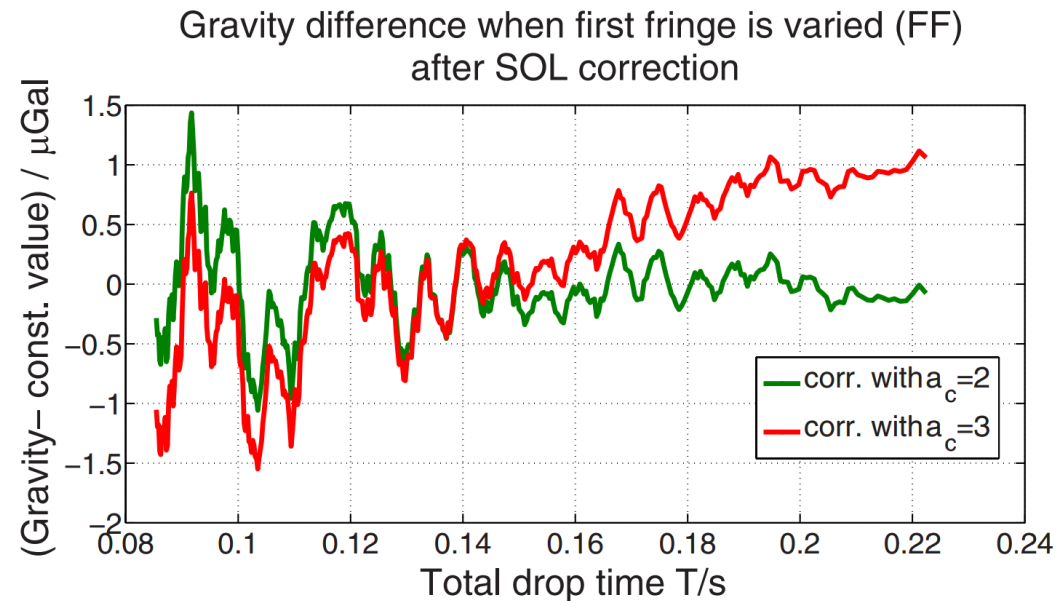
AGDAS, experiment

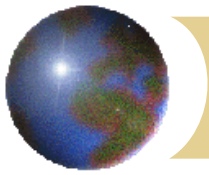
Correction due to finite speed of light

Surprising theoretical (Rothleitner and Francis 2011) and experimental results (Rothleitner et al. 2014), which should lead to correct FG5 (and gravity reference too) for 4 Gal.

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AGDAS is easy to use for detection if measured „g“ is dependent on the velocity of the test mass





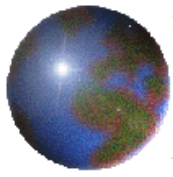
Summary and conclusions

SGNoise is providing useful information related to the quality of SG data. Repeated absolute measurements at the station might be immediately compared with gravity residuals.

There is a possibility to extend the noise level analysis to the sub-seismic and tidal bands.

The next work will be oriented for improvement of automatic data processing.

AGDAS is a powerful tool for advanced analysis of AG data as verification of the correction due to finite speed of light. Further, AGDAS has some advantages with respect to the „g“ outputs as 1) analysis of residuals for a given campaign, 2) accurate referencing of the results.



Thank you for your attention!

